1 NAPS: NAncay Preprocessing Software (written by Patrice Renaud)

NRT autocorrelator data reduction, calibration, exportation in SIR and FITS formats. (Document prepared by Jean-Michel Martin & Pierre Colom (V2 2008/03/26 - LaTeX 26-MAR-2008) This document is based on the on-line NAPS help file, written by P. Renaud. A manual (in french) is located in the folder /home/renaud/naps/works :

naps_userguide.doc (Feb. 2007 version)
naps_userguide.pdf (" ")

1.1 Definitions

 SCAN : raw observation of a source. Data written in nnnXXXXX.UIC files, with : nnn backend type (nco, fil...)
 XXXXX SCAN NUMBER
 UIC TELESCOPE USER ID

A SCAN includes some headers and one or several CYCLEs.

- CYCLE : Includes several PHASEs. A classical observation in position-switch mode would involve four phases :
 - one calibration phase (on source),
 - one ON source acquisition phase,
 - one calibration phase (off source),
 - one OFF source acquisition phase.

Some peculiar telescope setup files may involve more phases.

Each PHASE is made of several INTEGRATIONS (1 2 4 8 16 or 32 sec).

Each INTEGRATION contains 2 to 8 spectra, as the receiver may observe up to 4 frequencies, 1 to 4 polarization parameters

1.1.1 What is NAPS doing ?

Shortly, once a SCAN is selected, the software :

- eliminates bad integrations (in the acquisition phases) if required
- averages all integrations for each acquisition phase
- performs the phase arithmetic (for instance ON-OFF/OFF) for each cycle
- performs the flux density scale calibration
- averages the cycles

- displays and saves the result on disk (if required) in FITS format or in the NRT SIR data processing specific format.

1.2 Commands summary (with examples)

(NAPS commands may be typed in upper case or lower case characters.) When entering in NAPS, your .configdep file is read; see Miscellaneous section.

1.2.1 SCAN (observation number == raw data files) selection

| LIDI DOAN | LISUS UNE AVAILADIE SCAN NUMBERS |
|---------------------|--|
| SET DUIC | Selects the user ID of the scans to be processed |
| | SET DUIC 171 |
| SET SCAN | Selects the scan numbers to be processed |
| | SET SCAN 10210 11222-11320 |
| | SET SCAN * |
| SET SOU, SET (| COMP, SET REF selects the phase numbers for the |
| | ON, OFF and REFERENCE phases. |
| | (ON-OFF)/OFF is calculated for each cycle. |
| SET NOSOU, SET | NOCOMP, SET NOREF |
| | after SET NOREF, NAPS calculates only (ON-OFF) |
| | for each cycle. |
| | after SET NOSOU (and SET NOREF), the OFF phase is |
| | kept for each cycle. |
| SHOW SESSION | Shows session parameters |
| SET DIRDATA, SE | T DIRSIR, SET DIRFITS selects working directories |
| | SET DIRDATA /data1/guest3 |
| | SET DIRDATA /data2/calib (for raw test or cal. data) |
| SHOW SESSION | Shows session parameters |
| | |

1.2.2 Data visualisation (1D)

| SHOW CYCLE | Displays one cycle (from one to all correlator banks) |
|--------------|---|
| | SHOW CYCLE 1 |
| | SHOW CYCLE 1-10 (use NEXT to display the cycles 2-10) |
| | SHOW CYCLE 2, BATT=3 (third correlator bank BATTerie |
| | in french) |
| NEXT | Displays the next cycle |
| SHOW SPECTRE | Displays the integrations (elementary spectra) |
| | SHSPEC 1 2 3 (shortcut) |
| SHOW FILTRE | Displays the broad band filter data |
| | SHOW FILTRE 1 (all filters, CYCLE 1) |
| | SHOW FILTRE 1,4 (filter number 4, CYCLE 1) |
| SET X | Sets the units for the abscissa axis |
| | Arguments are : CHAN (channels) |
| | FREQ (frequencies) |
| | VEL (velocity) |
| | |

| | SET X VEL |
|--------------------|---|
| SET YPR | Sets the Y Plot Range |
| | SET YPR 3 (range = +/- 3 sigmas) |
| ENA VISU, DISA | VISU enables or disables the graphic display |
| ENA WNEW | a new graphic window is created for each plot |
| DISA WNEW | disables the WNEW mode |
| SET PSYM | selects the symbol used for the plots |
| PLOT | writes the plot in a POSTSCRIPT file |
| | |

1.2.3 2D Data display and processing

(2D time frequency display of each correlator bank.)

| DISP MODE=1 | Displays all cal. and acquisition integrations |
|--------------|---|
| DISP MODE=2 | Displays the acquisition integrations and marks the |
| | channels which show radio interferences |
| DISP MODE=3 | Displays the acquisition integrations using an ON-OFF |
| | algorithm (cleans the 'OFF' which is substracted to |
| | each integration) |
| DISP MODE=4 | Same, displays a cleaned ON-OFF spectrum and the percentage |
| | of removed data for each channel. |
| DISP MODE=5 | Displays all cycles after ON-OFF arithmetics. |
| | DISP MODE=1, BATT=1 |
| ENA/DISA CLR | Enables/disables the RFI detection algorithm |
| INTE2D | Same as command INTE, with RFI cleaning algorithm working |
| | in each cycle |
| SAVE2D | Same as command SAVE, with RFI cleaning algorithm working |
| | in each cycle |
| SET YPR | Curve min and max fixed, below each 2D image |
| | |

1.2.4 SCANs process

| PROC | SCAN | | Scan | proce | essir | ıg | (Tsys | are sh | own at | t the | e end) | 1 |
|-----------|--------|-------|-------|--------|-------|------|-------|---------|--------|-------|--------|----------------------------|
| | | | CI | LIP, N | WIN a | and | ILR | constra | ints a | are i | ised | |
| DUMP | MAP, | DUMP | SOU, | DUMP | SCAN | ١, ١ | DUMP | TABCAL, | LIST | TABO | CAL | |
| DISA, | /ENA E | BREAK | Save | all t | the a | sel | ected | SCANs, | with | one | SAVE | $\operatorname{command}$. |
| | | | ENA I | BREAK | (def | fau | lt) | | | | | |

1.2.5 USING a MASK for estimating the TSYS

| SET MASK | SET MASK 100-200, BATT=1-4 |
|-----------------|---|
| | channels 100 to 200, in spectra type 1 to 4 |
| ENA MASK | Enables MASK, necessary after MASK definition |
| DISA MASK | Disables MASK |
| SET DEFMASK | Back to default mask: 5 first and last points |
| SET NOMASK | Deletes the default mask |

1.2.6 Suppress cycles/banks

| SUP CYCLE | (shortcut SUPC) suppress cycles |
|---------------|---|
| | SUP CYCLE 2-3 4, batt=4 |
| | SUPC 2-3 4,0,0,4 |
| SET CLIP | Sets a limit for the cycles'Tsys. |
| | SET CLIP 100 (for all banks) |
| | SET CLIP 100 150 100 200 (for each bank) |
| ENA/DISA CLIP | Enables/disables the CLIPping |
| | DISA CLIP (default) |
| SET WIN | Sets a Tsys window, in order to reject cycles |
| | in units of rms (rms of the cycle's Tsys) |
| | SET WIN -2 2 |
| | The CLIP action is done before the WIN rejection. |
| ENA/DISA WIN | Enables/disables the WINdow-based selection. |
| | DISA WIN (default) |
| | |

1.2.7 Automatic removal of integrations

| (use | with care : t | this algorithm deletes always a certain percentage of | | | | | |
|------|--------------------------------------|---|--|--|--|--|--|
| inte | integrations in every phases/cycles) | | | | | | |
| S | ET ILR | Sets the ILR (Integration Limit Rms) | | | | | |
| | | SET ILR 1.2 | | | | | |
| E | NA/DISA ILR | Enables/disables this mode. 'Worst integrations' | | | | | |
| | | removal is performed by the command PROC SCAN | | | | | |
| | | DISA ILR (default) | | | | | |
| D | UMP SCAN | Displays the flag array for all the integrations | | | | | |
| R | ESET | Resets flagged integrations | | | | | |
| R | ESET ALL | Resets all cycles | | | | | |
| | | | | | | | |

1.2.8 Automatic removal of 'bad' channels

| (see | also | the | DISP | , IN | TE2D | and | SAVE2D | С | ommands | 3) |
|------|--------|-------|------|-------|-------|-------|----------|----|---------|------|
| SE | T CLR | ł | | Sets | the | CLR | (Channe | 91 | Limit | Rms) |
| EN | IA/DIS | SA CI | LR : | Enab | les/d | disab | oles thi | s | mode. | |
| | | | | ENA (| CLR | (defa | ult) | | | |

1.2.9 Integration and creation of final spectra (result files)

| INTE | Averages cycles (possibly after PROC SCAN) + display |
|----------------|---|
| INTE2D | Averages cycles using the 2D cleaning algorithm + display |
| SAVE | INTE + creation of a disk file (FITS or SIR format) |
| SAVE2D | INTE2D + creation of a disk file (FITS or SIR format) |
| SAVE CYCLE | Saves each cycle on disk (FITS or SIR format) |
| SAVE INTE | Saves each integration on disk (FITS or SIR format) |
| | |

(Be careful : you may create hundreds of disk files!)
-- GO Next scan

1.2.10 Miscellaneous

NAPS can work with script files. The script file names must include the extension .naps. -- CALL myscriptfile -- SET IDENT Changes the default NAPS identification of the final result file (for SIR format outputs only) Enables/disables the K-Jy calibration. -- ENA/DISA CAL -- ENA/DISA EFFI Takes into account (default) or not the radio telescope efficiency variation on declination in the calibration calculus. -- ENA/DISA TABCAL Prints the attenuation and Noise Diode values used by PROC SCAN -- ENA/DISA RHO Prints the rho(0) (first point of the autocorrelation function) values used by PROC SCAN. -- DUMP FREQ Prints the receiver's frequency setups. -- ENA/DISA FREQ Writes in addition the spectra with FREQuency axis (usefull for SIR format outputs) in the same SIR result files. Example of a .configdep file. (one line only !)

```
/data2/171/2008A/ /home/colom/sir/resultat/def/
/home/colom/sir/sorties/ /ps /ps /home/colom/fitsdata/
```

This file can be edited with a standard editor, or with the FIP tool.

1.3 More explanations Start of a Reference Manual

LIST SCAN Lists the scan numbers which are in the selected folder. This folder may be selected via the file .configdep, located in the user's home directory, or with the command SET DIRDATA. The file .configdep is created by the FIP software, and can be modified by the user.

Example for user dupont with the data on /data2/uic/2008A/ (holds on one line):

/data2/uic/2008A/ /home/dupont/sir/resultat/groupe1/ /home/dupont/sir/sorties/ /ps /ps /home/dupont/fitsdata/

where uic is the id 3 digits number of user dupont, and 2008A the semester of the observations. Do not forget the last / sign in the definitions.

SET DUIC Selects the user ID of the scan to process. The default DUIC is the standard UNIX/NRT user ID.

SET SCAN Selects the scan numbers to be processed. The character * selects all the scans available in the raw data folder. Example :

NAPS> SET SCAN 1435 10-20

The scan number 1435 and the scans with numbers between 10 and 20 will be processed successively. The first valid scan is selected and the system temperatures of the scan are typed.

NAPS applies a calibration (Jy) on the autocorrelator *and* the filter data, using calibration data obtained from systematic observations of series of standard radio sources.

Since November 13th 2000, the noise diode calibration sequences are recorded with both the autocorrelator and the filter bank. Each noise diode sequence is stored as three autocorrelator integrations of 1 second each (and three filter bank data points as well). With these integrations, it is possible to calibrate all polarisation setups, including the cross-polar ones like EW, EW^{*}, using the $\rho(0)$ power which is written after each spectrum.